

Institut Laue Langevin, Grenoble, France.

***Poor qubits make for rich physics:
noise-induced Zeno effect
& noise-induced Berry phases.***

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ICNF - Pisa, June 2009

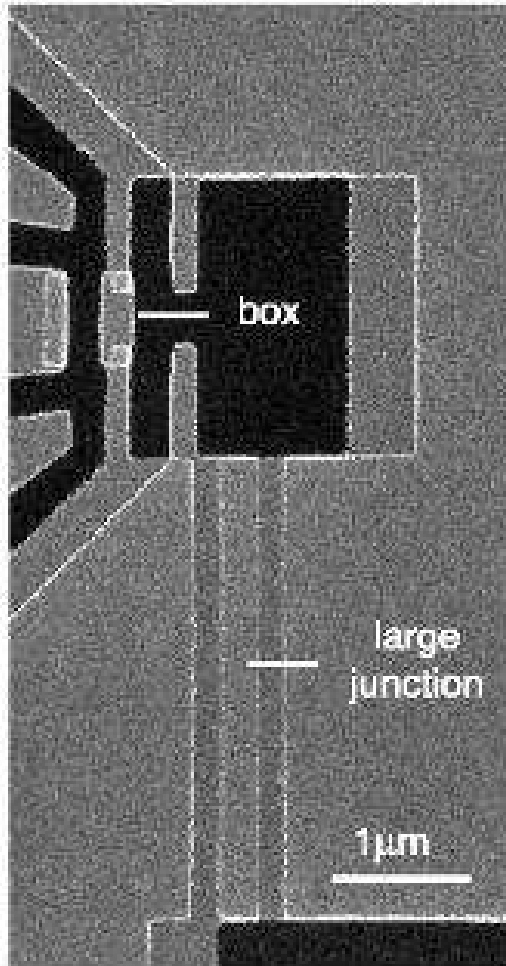
Summary

- ♣ Qubits in Markovian & non-Markovian noise
 - ◇ Intro: Quantum Zeno paradox (no noise)
 - ♣ Noise-induced Zeno paradox and "super"-Zeno paradox
 - ♣ Experts: dangers of adiabatic renormalization group

 - ◇ Intro: Berry phase (no noise)
 - ♣ Noise-induced Berry phase
-

Qubits : fully controllable two-level systems

Saclay qubit: Vion et al (2002)



$$\mathcal{H} = B_x \hat{\sigma}_x + B_y \hat{\sigma}_y + B_z \hat{\sigma}_z$$

+ NOISE

or quantum environment

Noise making quantum physics richer?

♣ NO NOISE :

System characterized by **wavefunction**, $|\psi\rangle = u|\uparrow\rangle + v|\downarrow\rangle$

two-levels \Rightarrow **TWO** independent variables

n -levels $\Rightarrow (2n - 2)$ independent variables

since $|u|^2 + |v|^2 = 1$ & drop overall phase

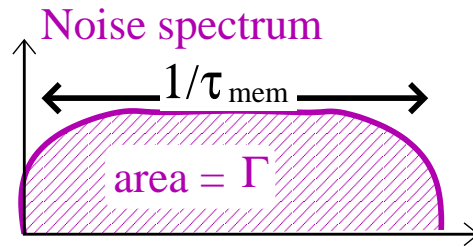
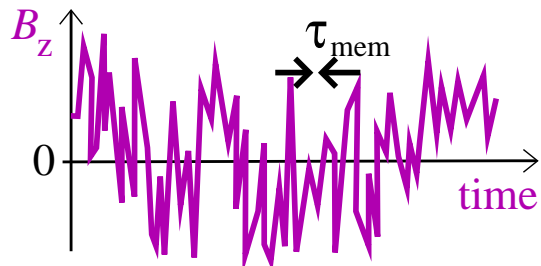
♣ WITH NOISE \equiv WITH QUANTUM ENVIRONMENT

System characterized by **density matrix**, $\rho = \begin{pmatrix} a & b + ic \\ b - ic & 1 - a \end{pmatrix}$

two-levels \Rightarrow **THREE** independent variables

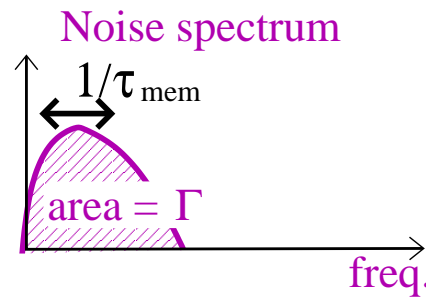
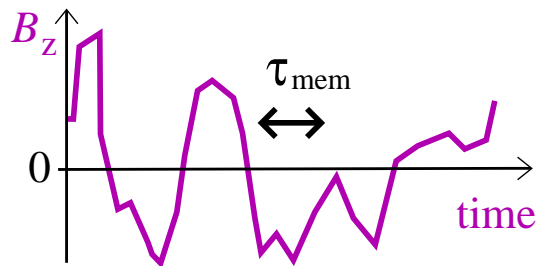
n -levels $\Rightarrow (n^2 - 1)$ independent variables

Markovian and non-Markovian noise

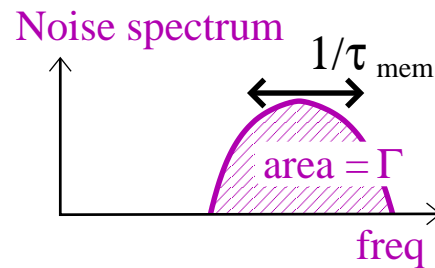
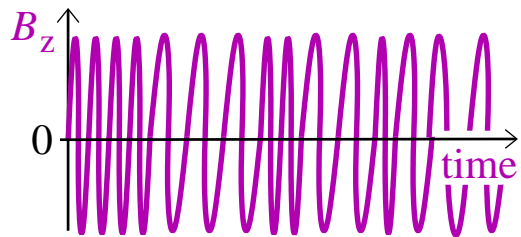


QUBIT (spin-half) +noise

$$\mathcal{H} = \mathcal{H}_{\text{qubit}} + \hat{\sigma}_z B_z(t)$$

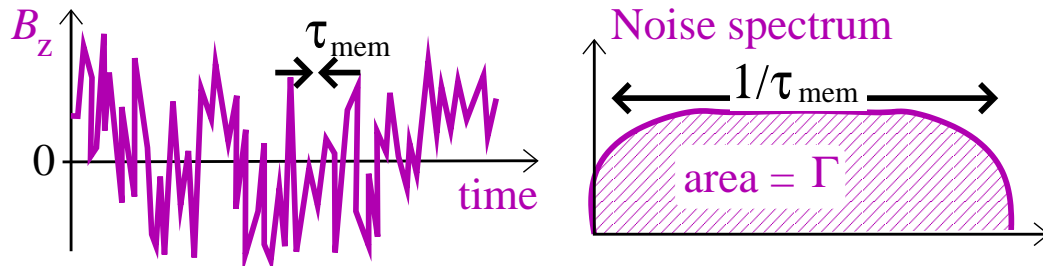


MARKOVIAN: $\Gamma \tau_{\text{mem}} \ll 1$
 \Rightarrow system weakly affected by
 noise on timescales τ_{mem}



NON-MARKOVIAN:
 $\Gamma \tau_{\text{mem}} \gg 1$

Markovian and non-Markovian noise



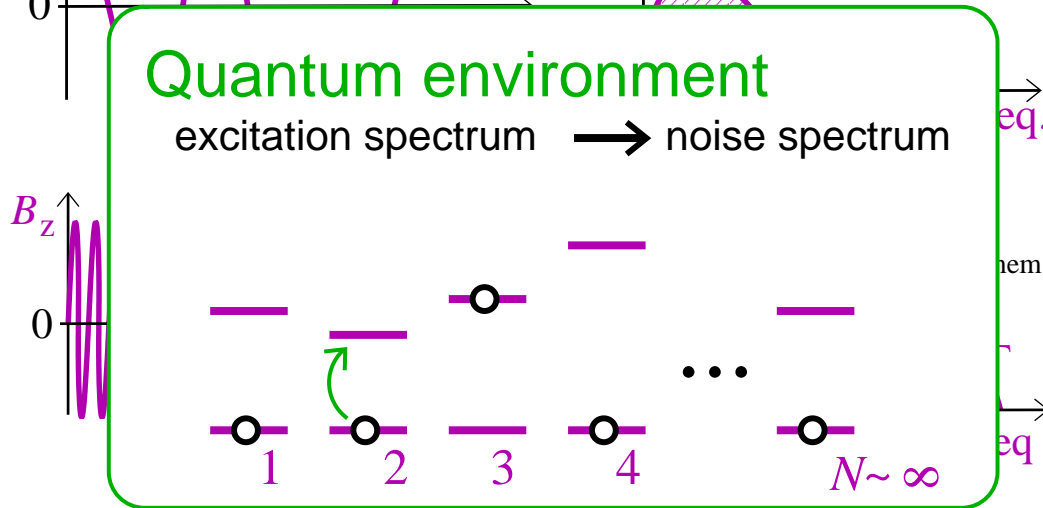
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NON-MARKOVIAN:
 $\Gamma \tau_{\text{mem}} \gg 1$



QUBIT (spin-half) +environ.

$$\mathcal{H} = \mathcal{H}_{\text{qubit}} + \hat{\sigma}_z K(\hat{a}^\dagger + \hat{a}) + \mathcal{H}_{\text{env}}$$

Quantum Zeno paradox

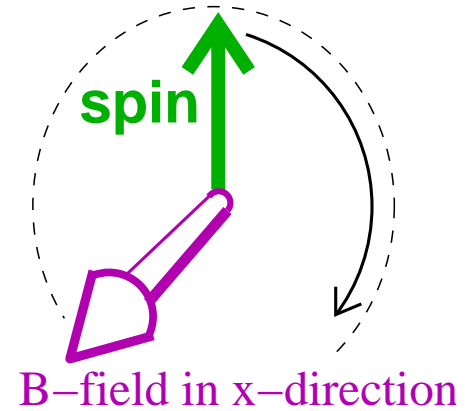
Misra- Sudarshan (1977)

~~Zeno (400BC) :~~

~~paradoxes with summing infinite series,
fully resolved by Cauchy in 1800s~~

QUANTUM clock:

spin-half precessing



OBSERVATION: measure spin along z-axis

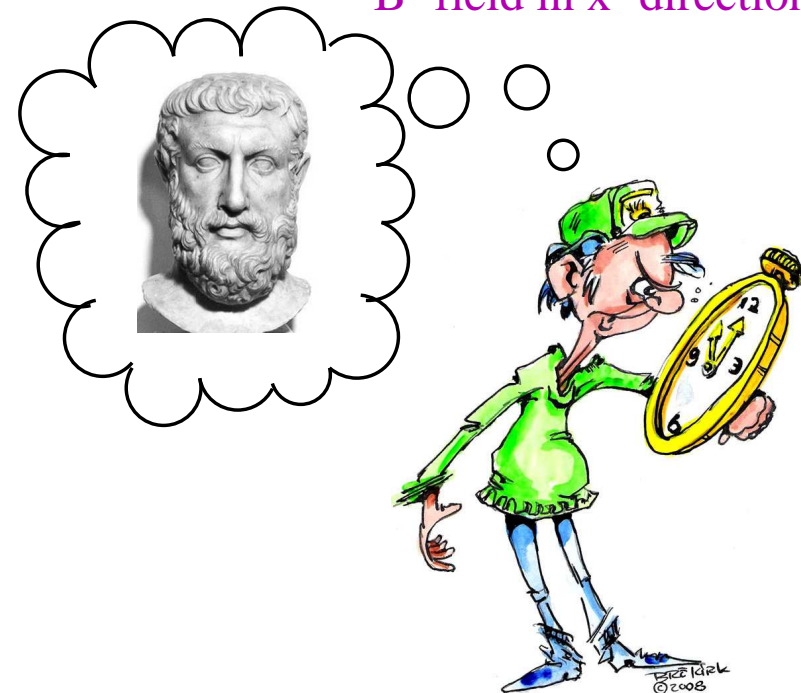
$$u|\uparrow\rangle + v|\downarrow\rangle \implies \begin{cases} |\uparrow\rangle \text{ with prob.} = |u|^2 \\ |\downarrow\rangle \text{ with prob.} = |v|^2 \end{cases}$$

Observation at intervals τ_{obs}

$$\longrightarrow \text{spin-flip time} \sim 1/(B_x^2 \tau_{\text{obs}})$$

cf. no observations

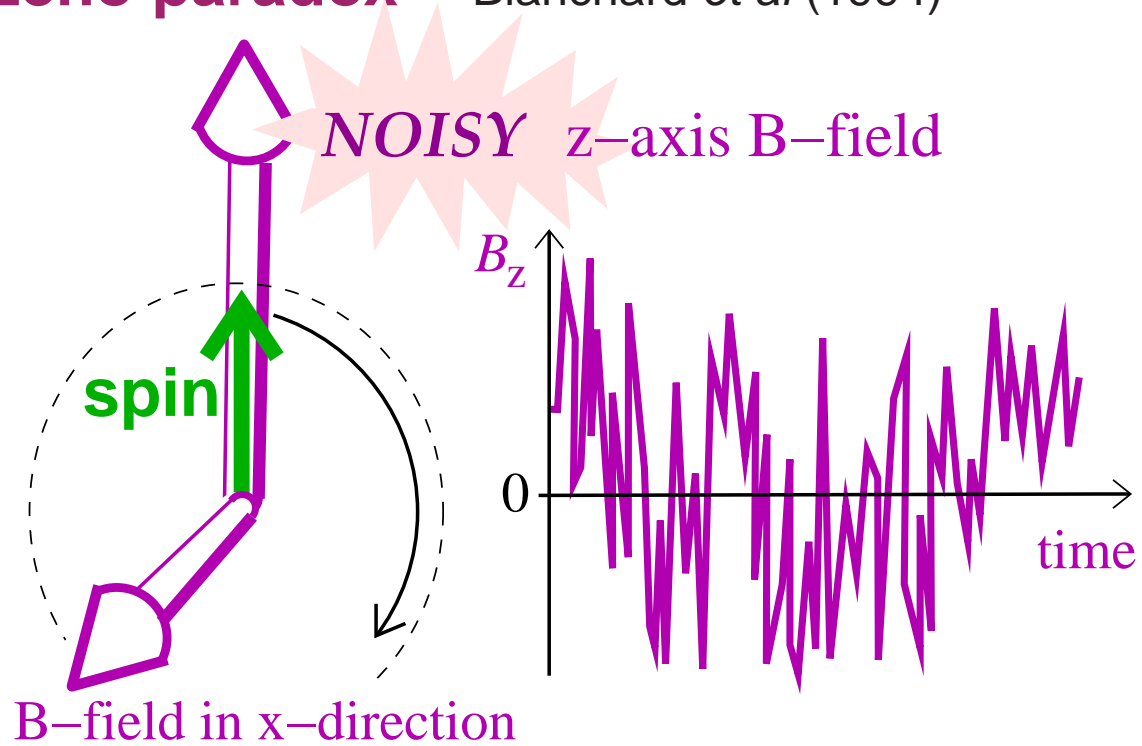
$$\longrightarrow \text{spin-flip time} \sim 1/B_x$$



Noise-induced Zeno paradox Blanchard et al (1994)

White-noise
— strictly Markovian

Damped harmonic oscillator eqn
for z-axis
spin polarization, s_z



underdamped ($B_x \gg \Gamma$) : $s_z \sim \exp[iB_x t - \Gamma t]$

OVERdamped ($B_x \ll \Gamma$) : $s_z \sim \exp\left[-\frac{B_x^2 t}{\Gamma}\right] \Rightarrow \text{spin-flip time} \sim \frac{\Gamma}{B_x^2}$

— analogy to usual quantum Zeno with $\Gamma \sim 1/\tau_{\text{obs}}$

Berry (1995)

Adiabatic renormalization

Legget et al (1987)

QUBIT +environment : $\mathcal{H} = B_x \hat{\sigma}_x + \hat{\sigma}_z K (\hat{a}^\dagger + \hat{a}) + \mathcal{H}_{\text{env}}$

qubit = SLOW

env modes = FAST

Born-Oppenheimer

no excited env modes

$$\langle j; A | j; B \rangle = 1 - K_j^2 / \Omega_j^2$$

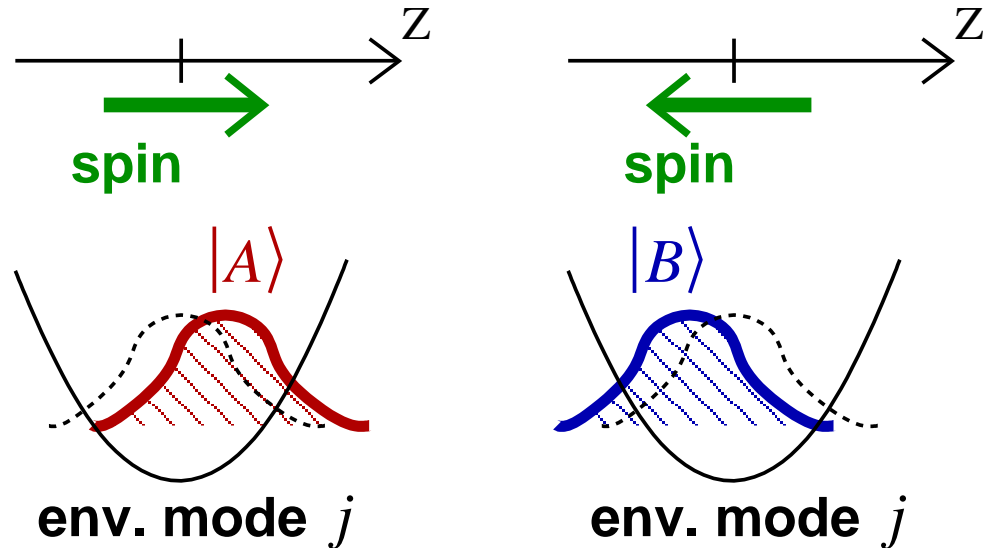
Franck-Condon

cf. Anderson orthogonality catastrophe

Overlap = $\prod_j \langle j; A | j; B \rangle = \exp \left[- \sum_j K_j^2 / \Omega_j^2 \right] \sim \exp[-\Gamma \tau_{\text{mem}}]$

$\hat{\sigma}_x \rightarrow \hat{\sigma}_x \times \exp[-F]$

spin-flip time $\sim \frac{\exp[+\Gamma \tau_{\text{mem}}]}{B_x}$



Exponentially slow \rightarrow “super” Zeno effect

For experts : beyond Born-Oppenheimer

♣ ADIABATIC RENORMALIZATION ARGUMENT

qubit + N modes = qubit + highest mode + (N-1) modes
= renormalized qubit + (N-1) modes

Use Born-Oppenheimer mode-by-mode
requires $B_x \rightarrow 0$ faster than Ω_{highest}
...but do “irrelevant” terms flow $\rightarrow 0$?



♣ POLARON MAPPING [Mahan's book, Aslangul et al 86, Dekker 87]

Magic mapping: non-Markovian \rightarrow Markovian for most env
+ perturbation theory on new \mathcal{H} (equiv to NIBA)

Terms missed by RG; in x,y-polarization but not z-polarization.

New rate $\sim \frac{B_x^2 B_z \tau_{\text{mem}}^2}{(\Gamma \tau_{\text{mem}})^{3/2}}$ powerlaw in Γ not exponential

Quick Intro to the Berry phase Berry [1984]

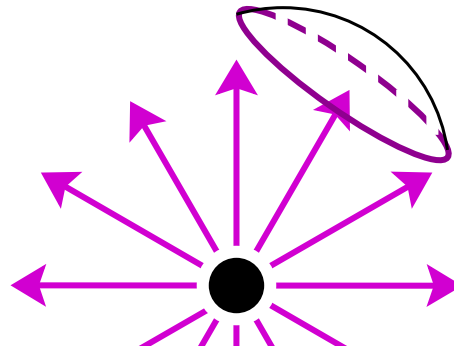
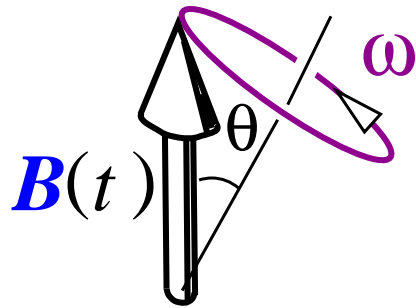
Observed in qubits Leek et al (2007)

Proposed uses: metrology

Pekola et al (1999)

quantum computing

Jones et al (2000)



Rotate B-field

$$\Phi_{\text{Berry}} = \text{solid-angle} = \text{monopole field} \\ = 2\pi(1 - \cos \theta)$$

But Berry phase not alone:

$$\Phi_{\text{total}} = \Phi_{\text{dyn}} + \Phi_{\text{Berry}} + \Phi_{\text{NA}}^{(1)} + \Phi_{\text{NA}}^{(2)} + \dots$$

where $\Phi_{\text{NA}}^{(\mu)} \sim (|B|t_p)^{-\mu}$

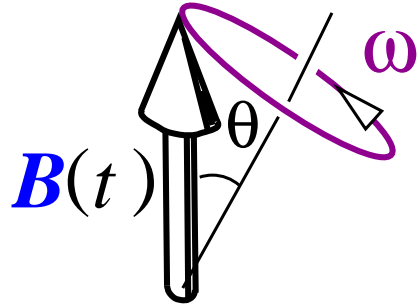
Measure Φ_{Berry} with accuracy of 1 in 1000 $\rightarrow Et_p \sim 10^{-3}$

Then $\Phi_{\text{dyn}} \sim 1000 \Phi_{\text{Berry}}$

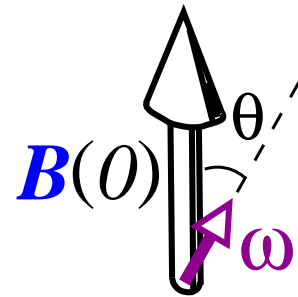
so must subtract Φ_{dyn} with accuracy of 1 in 10^6

Rotating frame for Berry phase and non-adiab phases Berry [1987]

Lab frame



Rotating frame



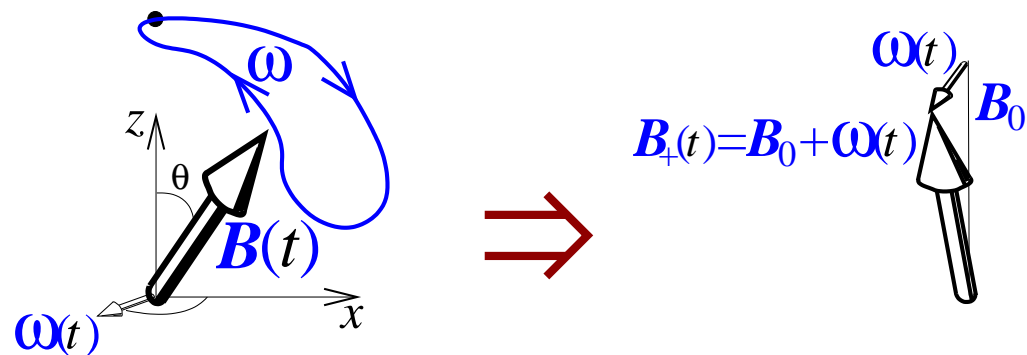
$$\Phi_{\text{total}} = |\mathbf{B} + \boldsymbol{\omega}|t_p = Bt_p + \omega t_p(1 - \cos \theta) + \mathcal{O}[\omega^2 t_p]$$

with $\omega t_p = 2\pi$

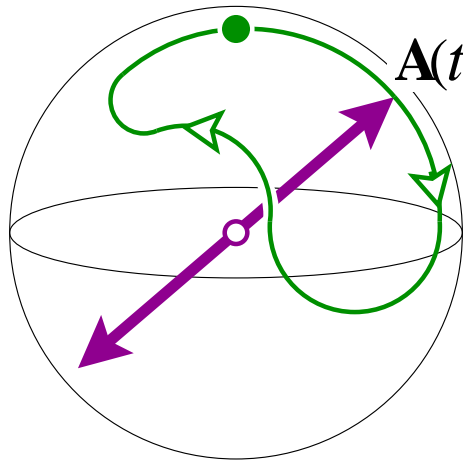
Go from strongly time-dependent problem
to (almost) time-independent problem

General transformation:

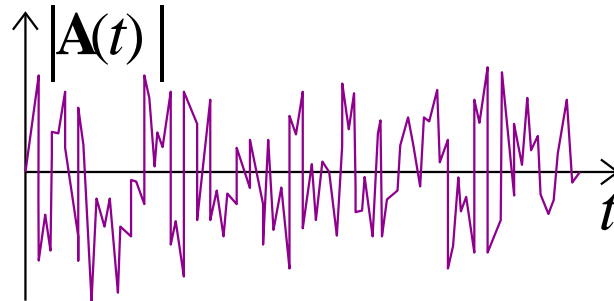
$$\mathcal{H}_{\text{rot}} = i[d\mathcal{U}/dt]\mathcal{U}^{-1} + \mathcal{U}\mathcal{H}_{\text{lab}}\mathcal{U}^{-1}$$



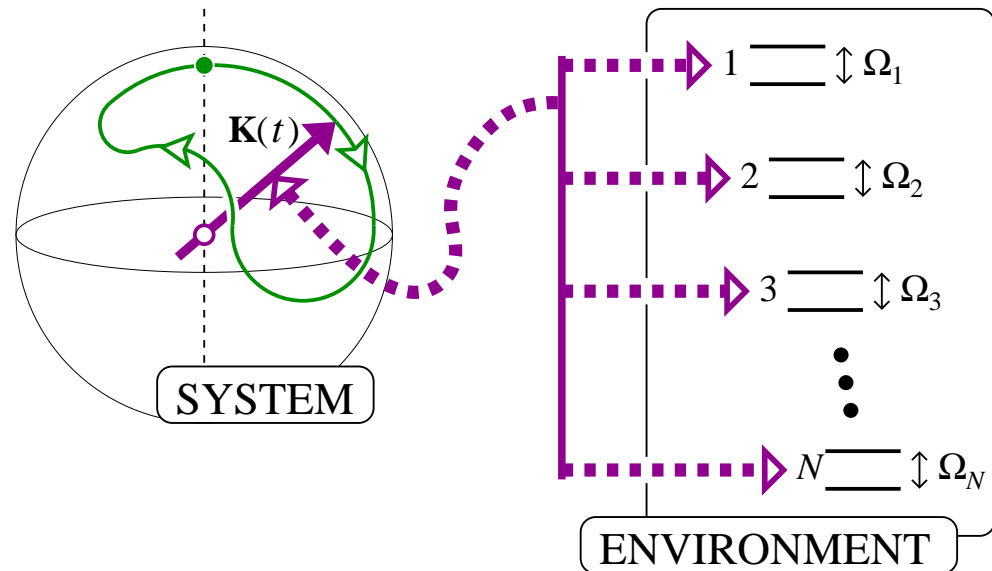
Making Berry phases with noise alone?



$$\mathbf{A}(t) = \mathbf{e}(t) \sum_{i=1}^N A_i \cos(\Omega_i t + \phi_i)$$



Quantum equivalent
(harder to realize)



Noise-induced Berry phase

♣ Rotate MARKOVIAN environment:

Carollo et al (2006) Dasgupta-Lidar (2007) Syzranov-Makhlin (2008)

Total phase **as for** conventional Berry phase

+ decoherence time $T_2 \ll t_P$

tricks to keep phase (decoherence-free subspace)

♣ Rotate NON-MARKOVIAN environment:

$$\Phi_{\text{total}} = \Phi_{\text{Berry}} + \Phi_{\text{NA}}^{(2)} + \dots \quad \text{with } \Phi_{\text{NA}}^{(2)} \sim 1/(\Gamma^{3/2} \tau_{\text{mem}}^{-1/2} t_p^2)$$

Measure Φ_{Berry} to 1 in 1000 $\rightarrow E_2 t_p \sim 10^{-3/2} \simeq 31$

Also decoherence not significant for $\Gamma t_p > 1$

less decoherence in long expt!

Why is noise-induced Berry phase like this?

Why a Berry phase?

Rotating frame:

“super”-Zeno effect suppresses ω_{\perp}

$\Rightarrow \omega_{\parallel}$ gives Φ_{BP}

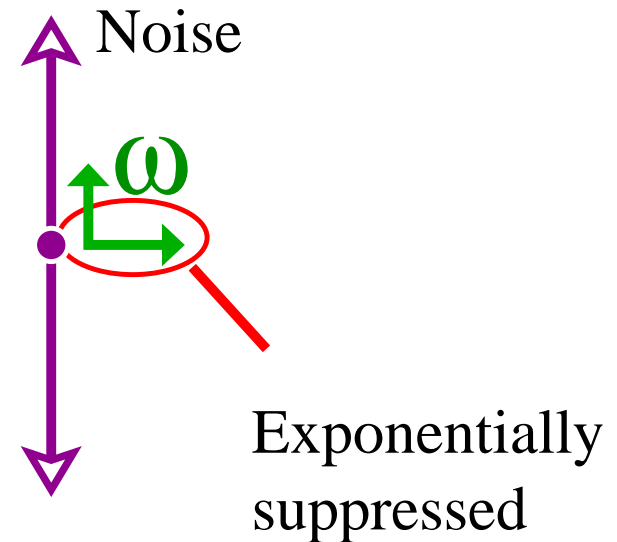
Why so little decoherence?

Magic of non-Markovian env.

No decoherence without transverse field

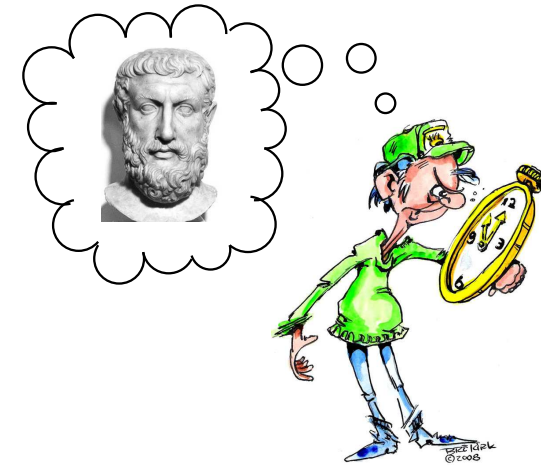
Why so little non-adiabaticity?

Not sure yet!



Conclusions

- ◇ Rich quantum physics with non-Markovian noise
- ◇ orthogonality catastrophe
or “super” Zeno effect



- ♣ non-Markovian noise makes a “better” Berry phase
get rid of unwanted Φ_{dyn} or $\Phi_{\text{NA}}^{(1)}$

- ♣ What other new things can noise do?
- ♣ Don't slip up

